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Publication details:

Climate Dynamics v. 41 Chapter No. 11-12 pp. 3073-3102 0930-7575 (ISSN)

Publication Date:

2013

Publisher DOI: http://dx.doi.org/10.1007/s00382-013-1676-1

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Climate Dynamics

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--Manuscript Draft--

Manuscript Number:	
Full Title:	The Indo-Australian monsoon and its relationship to ENSO and IOD in reanalysis data and the CMIP3/CMIP5 simulations
Article Type:	Original Article
Keywords:	Indian monsoon; Australian monsoon; Maritime Continent; ENSO; IOD; CMIP5; CMIP3; monsoon projection
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	Based on model fidelity in reproducing realistic monsoon characteristics and ENSO teleconnections, we objectively select 13 "best" models to analyze projections in the rcp8.5 scenario. Twelve of these models are from the CMIP5 ensemble. In India and Australia, most of these models produce 5 to 20% more monsoon rainfall over the 21st century than during the pre-industrial period. By contrast, there is no clear model consensus over the Maritime Continent.

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15 Abstract

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Based on model fidelity in reproducing realistic monsoon characteristics and ENSO teleconnections, we objectively select 13 "best" models to analyze projections in the rcp8.5 scenario. Twelve of these models are from the CMIP5 ensemble. In India and Australia, most of these models produce 5 to 20% more monsoon rainfall over the 21st century than during the pre-industrial period. By contrast, there is no clear model consensus over the Maritime Continent.

36

37 1 Introduction

The mechanisms that drive changes in the Indo-Pacific summer monsoon system are of con-38 siderable interest as this phenomenon affects many human activities and resources over broad 39 areas. The Indo-Australian monsoon consists of the Indian and South-East Asian summer 40 monsoon that occurs from June to September (JJAS), and of the monsoon that occurs in aus-41 tral summer (December to March, DJFM) over Australian and Maritime Continent (Neale 42 and Slingo, 2003). Contrary to popular understanding, the Australian and Maritime Conti-43 nent monsoon does not appear to be primarily driven by land-ocean temperature contrast 44 (Yano and McBride, 1998; Chao and Chen, 2001), and the importance of land-ocean contrast 45 for the Indian monsoon is still a matter of debate (Liu and Yanai, 2001; Chao and Chen, 46 2001). The presence of the Himalaya, however, plays a key role in the Indian Monsoon, 47 essentially by insulating warm moist air over India from cold dry air further North (Boos 48 and Kuang, 2010). 49

On interannual timescales, the India-averaged monsoon rainfall tends to be relatively 50 weak when it co-occurs with the development of an El Niño, and vice versa for La Niña. 51 Other sea surface temperature (SST) patterns such as the Arabian Sea upwelling (Izumo 52 et al., 2008) also seem to affect the regional distribution of monsoon within India on inter-53 annual timescales (Mishra et al., 2012). This picture is complicated by the Indian Ocean 54 Dipole (IOD, Saji et al., 1999) that modulates ENSO influence on the Indian summer mon-55 soon (Ashok et al., 2001; Ummenhofer et al., 2011). Australian monsoon rainfall also tends 56 to be weak during El Niños (Holland, 1986). The positive phase of the IOD (that peaks 57 in September-November, SON) also tends to weaken the following monsoon over the Aus-58 tralian/Maritime continent (Cai et al., 2005). 59

For the combined Indo-Australian monsoon system, Meehl (1997) and Meehl and Arblaster (2002) have described a tropospheric biennial oscillation (TBO) that seems to link the Indian and the Australian monsoons through ocean-atmosphere coupled mechanisms.

On longer timescales, the impact of climate change on the monsoon system is a major 63 concern. Climate change may directly affect the monsoon in two compensating ways: 1-64 warmer SSTs enable more evaporation and tend to increase the monsoon strength 2-SSTs 65 warm more in the equatorial region than in the Tropics, which tends to weaken the monsoon 66 circulation (Chung and Ramanathan, 2006; Krishnan et al., 2012). These two mechanisms 67 are tightly linked to possible change in SST global modes of variability, in particular ENSO 68 (El Niño Southern Oscillation) and the IOD (Indian Ocean Dipole) (Shi et al., 2008; Zhang 69 et al., 2012). Other factors may also influence the monsoon, as, for instance, the upper 70 tropospheric properties (Rajendran et al., 2012). 71

Over the last few years, the ability of general circulation models (GCMs) to realistically 72 simulate the Indo-Pacific monsoon and its teleconnections has been analyzed in the context 73 of the Coupled Model Intercomparison Program 3 (CMIP3), contributing to the Intergovern-74 mental Panel on Climate Change (IPCC) Fourth Assessment Report (IPCC, 2007). While 75 the link between ENSO and the Australian monsoon rainfall is rather well captured by the 76 CMIP3 models (Colman et al., 2011), the ENSO-rainfall relationship is poorly captured near 77 Papua-New Guinea (Cai et al., 2009). These authors have suggested that the ENSO-rainfall 78 relationship is affected by the so called "cold tongue bias" where SST is too cold along the 79 equator, and positive SST anomalies extend too far West during El Niño events (with a 80 significant impact on the Maritime Continent rainfall). In the CMIP3 ensemble, there is 81 no model consensus on how interannual variability of tropical Australian precipitation will 82

change in future climate (Moise et al., 2012). By contrast, a clear increase of future monsoon
rainfall has been found over the Maritime Continent (Smith et al., 2012). Finally, based on
the CMIP3 models, the South-East Asian summer monsoon is likely to undergo a slight
increase of precipitation in the future (IPCC, Meehl et al., 2007).

In this paper, we evaluate the Indo-Australian monsoon and its teleconnections to ENSO 87 and IOD in the CMIP simulations. We perform a combined analysis of simulations from 24 88 CMIP3 models and from 35 models taking part in the new Coupled Model Intercomparison 89 Program 5 (CMIP5). Results from 7 atmospheric reanalysis are also included as reanalysis 90 are often used as a proxy for dynamical observations (e.g. wind, pressure) to evaluate the 91 CMIP model dynamics, or to analyze mechanisms. Precipitation provides an integrated 92 assessment of the reanalysis skills (atmospheric model and data assimilation system) since 93 rainfall observations are generally not assimilated in the system (see section 2.2). Finally, 94 we select a subset of models that represent the Indo-Australian monsoon and its connections 95 to ENSO well, and we assess projected change in monsoon rainfall during the 21st century. 96

97 2 Datasets

98 2.1 Indices definition

To get an overview of the models skill, we use box-averaged indices rather than maps. We use two land-only monsoon rainfall indices for Australia and India (LAUS and LIND, Tab. 1, Fig. 1), as land-based rainfall has a direct influence on many human activities and resources, and because long-term rainfall data are only available over the land. Two other monsoon indices are also examined: AMAR and ISAS (Tab. 1, Fig. 1) that include rainfall over ocean and land over a larger domain. These indices are potentially better suited to examine
climate teleconnections and have been used previously to examine the TBO (e.g. Meehl and
Arblaster, 2002).

In addition, standard SST indices are used to describe the major tropical modes of 107 variability in the Indo-Pacific region (ENSO and IOD, see Tab.1 and Fig.1). It could be 108 argued that indices based on fixed locations may not fully capture the model dynamics 109 since simulated variability may have spatial biases compared to observations. However, it 110 is also important to capture modes of variability at realistic locations as this may affect the 111 propagation of teleconnection patterns to remote regions (Taschetto et al., 2009). For ENSO, 112 most of the CMIP3 models are not able to realistically reproduce distinct central Pacific El 113 Niño (also referred to as Modoki, or Warm Pool El Niño) and canonical El Niño (also referred 114 to as Cold Tongue El Niño) (Yu and Kim, 2010). Thus, CMIP3 models produce too much 115 coherence between NINO3, NINO34 and the El Niño Modoki Index (Cai et al., 2009). The 116 models skills concerning the representation of these two kinds of ENSO have improved in 117 CMIP5 (Kim and Yu, 2012; Taschetto et al., 2012). A majority of the CMIP3 and CMIP5 118 models still fail to capture the variance associated with these statistical modes realistically 119 (Roxy et al., 2012; Shamal et al. 2012, under preparation). For these reasons, we decide 120 to use the NINO34 index in this paper. It captures both kinds of ENSO without giving 121 too much importance to strong East Pacific (canonical) El Niño. The later have indeed 122 been suggested to have a weaker influence on the Indo-Australian monsoon than the central 123 Pacific El Niño (e.g. Taschetto and England, 2009; Kumar et al., 2006). For the IOD, Cai 124 et al. (2009) have shown that most of the CMIP3 models produce a SST dipole pattern that 125 is similar to observed, even though the amplitude of the cold tongue varies from model to 126

model (their Fig.10). Therefore, the simulated Indian Ocean Dipole Mode Index (DMI, as
defined by Saji and Yamagata, 2003) makes a reasonable index to represent the model IOD.
In this work, all the diagnostics related to the interannual variability of an index are made
after removal of the trend (linear least mean square fit) and of the climatological seasonal
cycle.

¹³² 2.2 Observation-based products and reanalysis

In this paper, we analyze precipitation data from 7 atmospheric reanalysis (lower part of Tab.1) and from gridded observational products (upper part of Tab.1). Some gridded observational products, such as GPCP, CMAP and TRMM-3B43, merge gauge analysis and satellite observations (available since the late 1970s). Before the satellite era, rainfall data are only available to any significant extent over land, and based on station measurements (GPCC, AWAP, APHRODITE). There are significant differences in the observation datasets, due to retrieval methods, treatments of uncertainties, and quality check (e.g. Yin et al., 2004).

Among the 7 atmospheric reanalysis, only ERAinterim uses a 4D-VAR data assimilation 140 scheme; the others use a 3D-VAR scheme. Rainfall from the various reanalysis is purely 141 model-generated (i.e. a forecast), since observed rainfall is not assimilated (see references in 142 Tab.1). An exception is MERRA whose atmospheric data assimilation has been developed 143 with a special focus on the hydrological cycle. While reanalysis generally show some skills in 144 reproducing the observed seasonal and interannual variability, their accuracy varies signifi-145 cantly across the regions (Bosilovich et al., 2008). Uncertainties in reanalyzed precipitation 146 may come from limitations in the dynamical models (e.g. convection, cloud microphysics, 147 complex topography), from uncertainties in the observations, and from the data assimilation 148

149 scheme itself.

The interannual variability of the various rainfall datasets for summer monsoon seasons 150 is compared to the observed variability in the Taylor Diagram in Fig. 2. Over Australia, 151 the AWAP dataset is generally considered as a reference, while the APHRODITE dataset 152 is chosen as a reference for South Asia and India because it has been developed with a 153 special focus on this region. These two datasets are based on weather stations, as is GPCC, 154 and cover a long time period. In general the observation and reanalysis datasets are more 155 consistent over Australia. The spread of reanalysis precipitation is larger for the Indian 156 monsoon, with outliers like NCEP-CFSR and, to a lower extent, NCEP-DOE-II. All the 157 observation products are correlated to AWAP by at least 0.95 in Australia. Correlation 158 coefficients between observations and APHRODITE are much lower over India, with, for 159 instance, CMAP being correlated to APHRODITE by 0.66. Note that 3B43 and 3B42 are 160 only weakly correlated to APHRODITE, but the overlap is only 10 years (their correlation 161 to GPCC is greater than 0.9 over the period 1998-2011). The correlation coefficient between 162 reanalysis and AWAP/APHRODITE is in the range 0.85-0.95 in Australia, and 0.35-0.85 in 163 India. It is possible that this difference in consistency between India and Australia could 164 be related to the Himalaya whose influence on the atmosphere is difficult to simulate, and 165 where in-situ observations are sparse and difficult to assimilate. A stronger influence of SST 166 for the Australian monsoon compared to the Indian monsoon may also improve consistency 167 in the reanalysis given that they are forced by observed SST. 168

Two SST datasets are used in this paper, HadISST and HadSST2 (Tab. 1). HadSST2 has a coarser (5-degree) resolution than HadISST, but no interpolation is used to fill grid points where observations are missing. As both datasets lead to similar results, only results from HadISST are shown here. For the sake of consistency, we also use SST from reanalysis when we produce diagnostics mixing SST and precipitation. It is important to keep in mind that the reanalysis use prescribed ocean SST, except for NCEP-CFSR which has a coupled ocean component.

176 2.3 CMIP3 and CMIP5 simulations

In this paper, we first analyze 24 CMIP3 simulations (Tab. 3) and 35 CMIP5 simulations 177 (Tab. 4) based on the historical simulations (called 20C3M in CMIP3 and historical in 178 CMIP5). The simulations start approximately in 1850 and end in approximately 2000 and 179 2005 for CMIP3 and CMIP5 respectively. The CMIP3 models and experiments have been 180 widely described in the literature over the last 5 years (e.g. Randall et al., 2007). Some 181 institutes have increased the resolution of their models from CMIP3 to CMIP5 (e.g. CNRM, 182 GISS, INMMRI). From CMIP3 to CMIP5 a large number of new experiments have been 183 included (Taylor et al., 2011). Some experiments now include a biogeochemical component 184 accounting for carbon cycles in the land, atmosphere, and ocean (Earth System Models, see 185 "ESM" in model names of Tab. 4). It should be noted, however, that the historical experiment 186 has prescribed gas concentrations (including CO_2). Some of the CMIP3 and CMIP5 models 187 have repeated historical (and future) experiments to form an ensemble with different initial 188 conditions (the initial state is taken in different points of the pre-Industrial simulation). 189

In section 3.3, we use a limited number of CMIP5 simulations to examine a future greenhouse gas and aerosols emission scenario. We use the representative concentration pathway rcp8.5 (Moss et al., 2010; Riahi et al., 2011). This scenario corresponds to a radiative forcing of approximately 8.5 W.m⁻² higher in 2100 than in the pre-industrial period. This is the ¹⁹⁴ most extreme scenario used to constrain the CMIP5 simulations in the sense that energy and ¹⁹⁵ industrial CO₂ emissions increase continuously until at least 2100 (whereas such emissions ¹⁹⁶ decrease from ~ 2080 in rcp6.0 and from ~ 2050 in rcp4.5).

Where we present information based on multi-model means, we first average across en-197 semble members of a given model, before averaging across the models. Where we consider 198 correlations between several model results, we assume that each model is different enough to 199 be considered independent (we thus probably over-estimate the significance since some mod-200 els are not strictly independent). A few of the CMIP3 models included water only (inmcm3-201 0) or water and heat (mri-cgcm2-3-2a, miub-echo-g, cgcm3 T47 and T63) flux adjustment. 202 Finally, some institutes have produced simulations from two models run at two different 203 resolutions (subscript LR/MR in model names of Tab. 3, 4), different cloud/convective pa-204 rameterization in the atmosphere model (e.g. IPSL-CM5A/IPSL-CM5B), or different ocean 205 models (e.g. GFDL-ESM2M/GFDL-ESM2G). In such cases, the two models are considered 206 separately, as independent models. Similarly, we consider that the CMIP3 version of a model 207 is independent from the CMIP5 version (e.g. gfdl-cm2-0/GFDL-CM3), and we even refer to 208 these two versions as "two models" in the following. 209

The acronyms used to refer to the institutes in Tab. 3 and Tab. 4 stand for the Environment Research & Technology Development Fund of the Ministry of the Environment, Japan (ERTDF), the Commonwealth Scientific and Industrial Research Organisation (CSIRO), the Bureau of Meteorology (BOM), the Beijing Climate Center (BCC) of the China Meteorological Administration (CMA), the Canadian Centre for Climate Modelling and Analysis (CCCMA), the Centro Euro-Mediterraneo per I Cambiamenti Climatici (CMCC), the National Center for Atmospheric Research (NCAR), the Centre National de Recherches Meteorologiques (CNRM) of Meteo-France, the European Centre for Research and Advanced Training in Scientific Computa-

tion (CERFACS), the Queensland Climate Change Centre of Excellence (QCCCE), the European Centre for 217 Medium-Range Weather Forecasts (ECMWF), the The National Key Laboratory of Numerical Modeling for 218 Atmospheric Sciences and Geophysical Fluid Dynamics (LASG), the Institute of Atmospheric Physics (IAP) 219 of the Chinese Academy of Sciences, the China Environmental Science and Sustainability Research Group 220 (CESS), the Tsinghua University (THU), the National Oceanic and Atmospheric Administration (NOAA), 221 the Geophysical Fluid Dynamics Laboratory (GFDL), the National Aeronautics and Space Administration 222 (NASA), the Goddard Institute for Space Studies (GISS), the Met Office Hadley Centre (MOHC), National 223 Institute of Meteorological Research (NIMR), the Korea Meteorological Administration (KMA), the Institute 224 for Numerical Mathematics in Moscow (INM), the Institut Pierre Simon Laplace (IPSL), the Atmosphere 225 and Ocean Research Institute (AORI) at the University of Tokyo, the National Institute for Environmental 226 Studies (NIES), the Japan Agency for Marine-Earth Science and Technology (JAMSTEC), the Max Planck 227 Institute for Meteorology (MPI-M), the Meteorological Research Institute (MRI), and the Norwegian Climate 228 Centre (NCC). 229

230 **3** Results

We first evaluate the mean summer monsoon rainfall, the amplitude of interannual variability, and the seasonal cycle in each model and reanalysis (section 3.1). Then, we assess the representation of the monsoon-ENSO and monsoon-IOD relationships (section 3.2). Based on these results, we select the most realistic models, and we show future projections of the monsoon (section 3.3).

3.1 Statistical properties of the historical Indo-Australian mon soon

The mean Indian and Australian summer monsoon rainfall is presented for each model and 238 reanalysis in Fig. 3-a. The spread in the observed mean summer monsoon rainfall is quite 239 large. For instance, the mean JJAS LIND is 6.8 mm/day in GPCC, versus 5.5 mm/day 240 in APHRODITE. Furthermore, the uncertainty in the mean rainfall in reanalysis is very 241 similar to the uncertainty in observations. We therefore choose to consider the multi obser-242 vation/reanalysis mean (black triangle in Fig. 3-a) as our reference here, with an uncertainty 243 envelope given by the two-dimensional PDF (Probability Density Function) of observations 244 and reanalysis (see caption of Fig. 3). The mean DJFM Australian monsoon rainfall based 245 on the multi-model mean is very similar to observational estimates in both the CMIP3 and 246 the CMIP5 models (triangles in Fig. 3-a). The mean JJAS Indian monsoon rainfall based 247 on the multi-model mean is under-estimated by $\sim 15\%$ in CMIP3, and by $\sim 19\%$ in CMIP5 248 (Fig. 3-a), though it lies within the 75% envelope of the observations/reanalysis. In both 249 CMIP3 and CMIP5 simulations, the relatively good skill of the multi-model mean hides a 250 wide spread in the mean monsoon rainfall, across individual CMIP3 and CMIP5 models: 251 from nearly no rainfall to twice as much rain as observed. The spread, as estimated by the 252 standard deviation, is 20% higher in the CMIP5 than the CMIP3 models for LIND, but 7%253 smaller for LAUS. Finally, there is a significant correlation between the average monsoon 254 rainfall in India, and that in Australia (r=0.56, p < 0.0001) which suggests that discrepancies 255 between models and observations are related to intrinsic model performance (e.g. convective 256 scheme, land surface scheme), not only to regional issues in the models. 257

²⁵⁸ The amplitude of the interannual variability is now evaluated through the standard de-

viation of summer-months-averaged rainfall, and presented in Fig. 3-b. The spread in the 259 observed values is larger here than for the mean, and has already been discussed for non-260 detrended time series in section 2.2. Due to this spread, we still consider the multi ob-261 servation/reanalysis mean as a reference, with an uncertainty envelope given by the PDF. 262 We nonetheless exclude two outliers from the multi observation/reanalysis mean and enve-263 lope calculation: NCEP-CFSR and NCEP-DOE-II (represented by μ), because these two 264 reanalysis present a much stronger interannual variability than any other reanalysis or ob-265 servation dataset. The standard deviation of both the Indian and the Australian monsoon 266 rainfall based on the multi-model mean is in remarkably good agreement with observations 267 in CMIP3 and CMIP5 (triangles in Fig. 3-b). This again hides a wide spread in the simu-268 lated amplitude of the interannual variability in both CMIP3 and CMIP5. The spread, as 269 measured by the standard deviation, is very similar in CMIP3 and CMIP5 for LIND, but 270 30% higher in CMIP3 than in CMIP5 for LAUS. Finally, there is a significant correlation 271 coefficient between the amplitude of the monsoon interannual variability in India and that in 272 Australia (r=0.52, p < 0.0001). This is probably related to the fact that there are common 273 drivers affecting the amplitude of the Indian and of the Australian monsoon (e.g. ENSO). 274 For both the Indian and the Australian monsoons, the correlation between the mean 275 and the interannual variability is relatively weak (r=0.12 for LIND and r=0.36 for LAUS). 276 This emphasizes the importance of evaluating a model both with regard to its mean and its 277 variability. For instance, the CMIP5 experiment from GFDL-ESM2M (represented by R) 278 has a realistic mean Australian monsoon rainfall, but its interannual variability is far too 279 strong.

As many modes of climate variability are phase-locked to the seasonal cycle, we also 281

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evaluate the later for each CMIP model in Fig. 4. By contrast to the mean and to the 282 interannual variability, the seasonal cycle is robust across the observations and reanalysis 283 (see the small RMS errors in Fig. 3-c and 4). Based on the multi-model mean seasonal cycle, 284 there is a clear improvement from the CMIP3 to the CMIP5 simulations (the RMS error 285 is reduced by 25-30%, see triangles in Fig. 3-c). However, the simulated seasonal cycle are 286 generally either too peaked in February (e.g. IPSL-CM5B-LR), or have an overly indistinct 287 monsoon season with high rainfall extending into April-May and October-November (e.g. 288 CCSM4). The maximum generally occurs on the right month, despite a few exceptions as 289 mpi-echam5 (Fig. 4). In the 20 CMIP simulations that best represent the seasonal cycles 290 (ranked by RMSE in Fig. 4), only 4 (3) are from CMIP3 for LAUS (LIND). It should also be 291 noted that the best CMIP3 model both for LIND and LAUS (mri-cgcm2-3-2a, represented 292 by t) has both heat and water flux corrections. Finally, there is also a significant correlation 293 coefficient between the amplitude of the RMSE of the simulated seasonal cycle in India and 294 that in Australia (r=0.55, p < 0.0001). 295

A first model selection is made, based on the three statistical properties of the Indo-296 Australian monsoon depicted in Fig. 3. As mentioned above, the mean summer monsoon 297 rainfall and its interannual variability show a significant spread in the observations. We 298 take this into account, and select the models that are within the contour enclosing 99.9%299 of the observations/reanalysis PDF integrative (blue contour in Fig. 3-a,b). This value is 300 found empirically, in such a way to keep a sufficient number of models in the selection pro-301 cess. We do the same selection with regards to the seasonal cycle (Fig. 3-c), except that 302 we extend the contour so that it encloses 99.999% of the observations/reanalysis PDF in-303 tegrative. This extension is needed because the spread in the observed/reanalyzed seasonal 304

cycle is very weak, and because the CMIP simulated seasonal cycles are significantly distinct 305 from the observed ones. Our overall method of selection allows eliminating models based 306 on the observations, taking uncertainty into account. Using the contour values mentioned 307 above leads to a selection of 19 models that are indicated in Tab. 5. Only one of these 19 308 models (gfdl-cm2-0) is from the CMIP3 ensemble. It should be noted that these models are 309 not entirely independent because some components are commonly used in several models. 310 For instance, CESM1-CAM5, CESM1-FASTCHEM, CCSM4, FIO-ESM, NorESM1-M, and 311 NorESM1-ME include an atmospheric component based on the NCAR Community Atmo-312 spheric Model (CAM), even though versions differ across the institutes. The Hadley Centre 313 atmospheric model is also the base of the atmospheric component in ACCESS1-0, Had-CM3, 314 and HadGEM2-AO. The models ACCESS1-0, GFDL-CM3, and gfdl-cm2-0 have an ocean 315 component based on the GFDL Modular Ocean Model (MOM). Finally, the Parallel Ocean 316 Program (POP), which originated from the same historical base as MOM in the 1990s, is 317 also a common base for the ocean component in CESM1-CAM5, CESM1-FASTCHEM, and 318 CCSM4. 319

320 **3.2** Relationship between SST modes and the Indo-Australian 321 monsoon

As mentioned in the Introduction, the effect of greenhouse gases and aerosols on the Indo-Pacific monsoon will be modulated by ENSO and the IOD. Therefore, it is essential to have a realistic representation of the monsoon-ENSO and monsoon-IOD relationships in the historical experiments. This is analyzed in the present sub-section, starting with the Australian monsoon-ENSO relationship. The TRMM observational products 3B42 and 3B43 ³²⁷ are not shown here, given the short record period.

328 Australian monsoon-ENSO relationship

Both the CMIP3 and the CMIP5 simulations show a moderate anti-correlation (-0.3) to 329 NINO34 (Fig. 5-a), i.e. an Australian monsoon occurring during an El Niño event tends 330 to be weaker than normal. This anti-correlation is significant at the 90% level for most 331 of the CMIP3 and CMIP5 models (see upper quartiles), and is in good agreement with 332 long observation timeseries (GPCC or AWAP, and HadISST, also already noted by Holland, 333 1986). Each model is shown separately in Fig. 6-a. As noted by Colman et al. (2011), a few 334 CMIP3 models do not show significant ENSO-monsoon relationship at zero lag (e.q. giss 335 models), while several CMIP3 simulations (e.g. csiro-mk3-5) produce too strong an anti-336 correlation at zero lag. This is still the case in the CMIP5 simulations, but correlations at 337 zero lag are generally closer to GPCC and AWAP values. Thus, 33% of the CMIP3 models 338 produce correlations in the range ± 0.1 of the observed correlation coefficient (-0.35), while 339 46% of the CMIP5 models do so. Surprisingly, this score is not better for the reanalysis, 340 since only 3 of 7 are in this ± 0.1 range (NCEP-CFSR, NCEP-NCAR-I, and MERRA). Cai 341 et al. (2009) demonstrated that in the CMIP3 simulations, models which produce an ENSO 342 with a strong interannual variability tend to have a strong rainfall-NINO34 anti-correlation 343 at zero lag. This is also what we find considering CMIP3 and CMIP5 models together (see 344 pink line and circles in Fig. 6-b). A few models, in particular from the CMIP3 ensemble, 345 show that a strong monsoon is followed, one year after, by an El Niño event (vice versa for La 346 Niña). This is probably generally attributed to the fact that some CMIP3 models produce 347 an ENSO with too strong a quasi-biennial component (e.g. gfdl-cm2-0 and miub-echo-g, 348 Fig. 6-b). 349

An interesting feature in Fig. 5-a is the anti-correlation peak (-0.15) that is found in 350 the multi observations/reanalysis mean 2 years prior to the monsoon, but which is totally 351 absent from the CMIP simulations. When considering AWAP only over 100 years (dashed 352 blue curve), this peak is reduced by half and far from significant at the 90% level. This 353 peak is nonetheless interesting because it is opposite as what is expected from ENSO, since 354 the later is anti-correlated with itself 1.5-2.5 years before (Fig. 5-d). As ENSO has been 355 suggested to have a strong inter-decadal to centennial variability (Wittenberg, 2009), we 356 raise the question as to whether such a peak can be captured by the models over a 30-357 year period. We extract the 30-year period from each 150-year simulation among the several 358 available ensemble members that gives the strongest anti-correlation at year -2. As shown in 359 Fig. 5-c, most of the CMIP models are able to simulate at least one 30-year period presenting 360 an anti-correlation peak similar to the observed one. We therefore suggest that this peak 361 is an artifact originating from the low frequency variability. Over such 30-year periods, 362 the probability of Australia to Australia successful TBO transition is slightly increased (not 363 shown), though the predictability associated with this transition remains non significant (not 364 shown, see also Li et al., 2012). 365

We now expand the region used to define the Australian monsoon, by including the Maritime Continent (land and ocean, see Fig. 1). The relationship between ENSO and monsoon rainfall is now stronger: the anti-correlation between AMAR and the concomitant NINO34 is -0.6 for the multi observations/reanalysis mean (Fig. 5-b), and even reaches -0.8 for CMAP and GPCP (Fig. 7-a). All the reanalysis have strong anti-correlations (moderately strong for NCEP-NCAR-I and ERA40), which is consistent with an oceanic control of the monsoon in this region as most of the reanalysis are forced by observed SST. It should nonetheless

be noted that shorter observational records are involved over the Maritime Continent (lim-373 ited to the satellite era). Such strong anti-correlations are generally not reached in CMIP 374 simulations, with a multi-model mean of -0.1 and -0.3 for the CMIP3 and the CMIP5 mod-375 els respectively (Fig. 5-b). The CMIP5 models thus have better skills in capturing the 376 AMAR-ENSO relationship, and more than 25% of CMIP5 models give anti-correlations that 377 reach the range of reanalysis (see lower quartiles in Fig. 5-b). Comparing Fig. 7-a and 7-b, 378 it is clear that most models that reproduce the observed AMAR-ENSO concomitant anti-379 correlation also simulate a realistic ENSO pattern along the equator, with a cold anomaly 380 in the climatological warm pool region (*i.e.* approximately at the location of the Maritime 381 Continent). This result is not surprising for this maritime region where convection is trig-382 gered by warm SSTs. Catto et al. (2012) have shown that only a few of the CMIP3 models 383 are able to simulate a realistic relationship between NINO34 and SST North of Australia. 384 Among their models, gfdl-cm2-0 has been shown to have the best skills. This is also what 385 we find in Fig. 7-b, if we exclude csiro-mk3-5 which produces an unrealistic ENSO pattern 386 as well as an unrealistic correlation between AMAR and ENSO 2 years in advance. Apart 387 from these two CMIP3 models, the 19 CMIP models giving the strongest AMAR–NINO34 388 anti-correlation (< -0.36) are from CMIP5. 389

³⁹⁰ Indian monsoon-ENSO relationship

We now evaluate the Indian monsoon-ENSO relationship. As shown in Fig. 8-a,b, the Indian and South-Asian monsoons tend to be weak when their development is concomitant with the development of an El Niño, and vice versa for La Niña. The correlation between the land-based Indian JJAS rainfall (LIND) and NINO34 reaches -0.60 in APHRODITE, and -0.45 in the multi observations/reanalysis mean. The same correlation but for ISAS instead

of LIND is slightly weaker (-0.35), and slightly less persistent (Fig. 8-b). The LIND-NINO34 396 correlation at zero lag, based on the multi-model mean, is underestimated by the CMIP5 397 models and even more by the CMIP3 models (Fig. 8-a). The ISAS-NINO34 correlation at 398 zero lag is realistic in the CMIP5 models, and slightly underestimated in the CMIP3 models 399 (Fig. 8-b). However, the most striking feature of CMIP correlation curves is that they present 400 a very gentle slope the year before the monsoon, as compared to a very steep slope in the 401 observations (Fig. 8-a,b). This is also true after the monsoon season for LIND. We link this 402 bias to the ENSO characteristics in the following paragraph. 403

In the observations and reanalysis, a positive (negative) anomaly of NINO34 in June-404 July, i.e. at the beginning of the Indian monsoon, is more related to a developing El Niño 405 (La Niña) event than to a terminating event. Indeed, the correlation between NINO-34 in 406 June-July (beginning of the Indian monsoon) and NINO34 in December prior to the mon-407 soon (December corresponds to the mature stage of ENSO) is only 0.15 (see black curve in 408 Fig. 8-c); by contrast, the correlation between NINO-34 in June-July and NINO34 in Decem-409 ber following the monsoon is high (0.80, see Fig. 8-c). This asymmetry is much less marked 410 in the CMIP simulations, in particular in CMIP5, where the equivalent correlation is 0.30 411 in December before the Indian monsoon, and 0.70 in the following December (green curve 412 in Fig. 8-c). In other words, the June-July NINO34 anomalies are mostly related to devel-413 oping El Niño or La Niña events in the observations, whereas they are also partly related to 414 the termination of previous-boreal-winter events in CMIP5 (and CMIP3 to a lower extent). 415 This can be explained by the too large spread in the seasonal cycle of NINO34 (and ENSO 416 in general) in the CMIP5 simulations, as described by Taschetto et al. (2012). During the 417 second half of the Indian monsoon (August-September), the NINO34 anomalies are good 418

precursors of developing El Niño or La Niña events (that will be mature in the following 419 December) both in the observations and in CMIP models. Indeed, the correlation in Fig. 8-d 420 is near zero in December prior to the monsoon, and 0.90 in the observations/reanalysis and 421 CMIP5 simulations (0.80 for CMIP3) in December after the Indian monsoon. To summarize, 422 as CMIP5 NINO34 June-July anomalies are too strongly related to previous-boreal-winter 423 El Niño or La Niña events, they produce too large negative correlations between the In-424 dian monsoon and the previous-boreal-winter NINO34 (Fig. 8-a,b). This bias is slightly 425 lower in the CMIP3 simulations, but the correlation between the Indian monsoon and the 426 following-boreal-winter NINO34 is better in CMIP5 than in CMIP3 (Fig. 8-a,b), due to the 427 better consistency between boreal-summer NINO34 anomalies and NINO34 anomalies in the 428 following boreal winter (Fig. 8-c,d). We have to note that the discussion here is from the 429 perspective of a periodic ENSO. There have been periods characterized by a string of El 430 Niño Modoki events during the early 1990s (Ashok et al., 2007). Our description does not 431 account for these particular periods, but holds on average. 432

Results for individual models are shown in Fig. 9. By looking at the symmetry of the negative correlation to each side of August at zero lag, it is possible to assess if a model produces monsoons mostly correlated to previous ENSO events (e.g. miroc3-2-hires, ACCESS-1.3, that sow strong asymmetric blue bar left of the year-0 line), partly correlated to previous ENSO events (e.g. FIO-ESM shows blue bar to either side of the year-0 line), or mostly correlated to developing ENSO events (e.g. HadGEM2-ES shows strong asymmetric blue bar right of the year-0 line).

440 Australian monsoon-IOD relationship

We next evaluate the relationship between the IOD and Australian rainfall in the various 441 datasets. There is a weak anti-correlation (hardly significant at the 90% level) between the 442 DJF land-based Australian monsoon (LAUS) and the previous SON (September to Novem-443 ber) Indian Dipole Mode index (DMI) for both AWAP and GPCC (represented by δ and γ 444 respectively in Fig. 10-a). This correlation is weaker and not significant for the reanalysis. 445 Both CMIP3 and CMIP5 multi-model means indicate that the models, in average, produce 446 an anti-correlation similar to that of AWAP and GPCC (-0.2). However, the spread across 447 models is large. Interestingly, there is a clear relationship between the strength of the previ-448 ous DMI-monsoon relationship, and the concomitant ENSO-monsoon relationship in CMIP 449 models (correlation of 0.73, see Fig. 10-a). Over the Maritime continent, models and reanal-450 ysis produce much stronger IOD-rainfall anti-correlation than with LAUS (Fig. 10-b), and 451 there is also a clear relationship between the strength of the previous DMI-monsoon rela-452 tionship, and the concomitant ENSO-monsoon relationship in CMIP models (correlation of 453 0.66). These results are consistent with the ENSO-IOD seasonal phase-locking relationship 454 that has been described in the literature (Annamalai et al., 2005; Behera et al., 2006; Luo 455 et al., 2010). We have also found that most of the CMIP models reproduce an Indian Ocean 456 Basin-wide Warming (IOBW, Lau and Nath, 2003; Chowdary and Gnanaseelan, 2007; Saji 457 et al., 2006; Ashok et al. under preparation 2012). The simulated IOBW-Australian mon-458 soon relationship is very similar to the ENSO-monsoon relationship shown in Fig. 5 (not 459 shown). However, as the IOBW is essentially a forced response to ENSO, it is not possible 460 to separate their relative contribution without specific idealized experiments (e.g. Taschetto 461 et al., 2011). 462

463 Model selection

In the previous subsection, we have selected a subset of models that have good overall 464 monsoon skills in the historical period (Tab. 5). We now examine to what extent these specific 465 models can reproduce the relationships to ENSO, in order to improve the model selection. 466 In the previous analysis, it appears that the AMAR-ENSO relationship is the most robust in 467 the observations (the strongest in terms of correlation). It is also an important area in terms 468 of North-South monsoon connection, at least on the biennial scale (Meehl, 1997); further, 469 there is a clear consensus among observations and reanalysis (Fig. 7). On this basis, we first 470 want to reproduce the link between ENSO SST patterns near the climatological warm pool 471 location and the convection over the Maritime Continent, i.e. we choose to exclude all the 472 models below the 99% significance level for the concomitant ENSO-AMAR correlation (i.e. 473 all the models below the horizontal dashed line in Fig. 7). Results are summarized in Tab. 5). 474 Apart from GFDL-CM3, the excluded models suffer from an unrealistic cold anomaly during 475 El Niño events (and vice versa during La Niña) in the Western Pacific. Then, we decide to 476 allow a bias of ± 0.3 in the DJF LAUS-NINO34 anti-correlation, i.e. we keep correlation 477 values between -0.15 (90% significance level) and -0.75 (i.e. between the dashed lines in 478 Fig. 6-a; see also Tab. 5). We also want some relationship between ENSO and the Indian 479 monsoon. Models with insignificant LIND-NINO34 correlation in JJAS at zero lag (Fig. 9) 480 are therefore excluded (Tab. 5). As the spread in the monsoon-IOD relationship is large in 481 both the observations and models, it is difficult to find clear outliers. Therefore, we do not 482 select the models on an IOD-monsoon relationship basis. Doing this, we select 13 models in 483 the 19 previously selected. Twelve of them are from CMIP5. 484

485 3.3 Future monsoon projections

Based on the subset of most realistic models, we now analyze the projected change of the Indo-Australian monsoon in the emission scenario rcp8.5 (see section 2.3). We do not analyze the simulations from the gfdl-cm2-0 CMIP3 model since the emission scenarios in CMIP3 were different from the rcp ones. Among the 12 selected CMIP5 models, all but CESM1-FASTCHEM were available for the rcp8.5 scenario at the time of writing, with several ensemble members for six of them (Tab. 5).

The evolution of the monsoon rainfall in the different boxes used in this paper is shown in Fig. 11. We average the indices over 50-year periods to increase the statistical significance. The confidence interval for each period is thus proportional to $s/\sqrt{50 N}$, where s is the interannual standard deviation and N the number of ensemble members (t-statistics, e.g. Von Storch and Zwiers, 2002). Two 50-year periods are considered significantly different if there is no overlap of the error bars in Fig. 11 (we consider a confidence interval at the 90% level).

Only two of the 11 models show a significant increase in monsoon rainfall over Australia 499 during the historical period (MIROC5 and FGOALS-s2, Fig. 11-a). This contrasts with 500 the results from Shi et al. (2008) and Smith (2004) who have reported an increase of the 501 observed land-based Australian monsoon rainfall in the 20th century. Now considering the 502 future projections, we find that nine of the 11 CMIP5 models show a significant rainfall 503 increase at the end of the 21st century as compared to the pre-industrial period. This increase 504 is in the range 12-22%. The two remaining model do not show a significant trend from 1850505 to 2100. 506

⁵⁰⁷ None of the selected CMIP5 models shows an increase in monsoon rainfall over the

Maritime Continent during the historical period (Fig. 11-b), but one model (HadGEM2-AO) 508 produces less monsoon rainfall at the end of the 20th century than during the pre-industrial 509 period. There is no clear consensus between the models concerning the future monsoon 510 rainfall over the Maritime Continent: three models produce less rainfall in 2050-2099 than 511 during the pre-industrial period (FGOALS-s2, HadGEM2-AO, CanESM2); two models show 512 trends that are not significant at the 90% level (ACCESS1-0 and CESM1-CAM5); the six 513 remaining models produce between 3 and 13% more monsoon rainfall at the end of the 514 21st century as compared to the end of the 19th century. 515

The picture is not clear either for the rainfall evolution over India and South-Asia over 516 the historical period (Fig. 11-c,d). Indeed, the majority of models do not show a significant 517 trend, while two models produce slightly more rainfall at the end of the 20th century, and 518 2-3 models produce less rain during this period. These results must be considered in the 519 perspective of Goswami et al. (2006)'s results: using observations, they have shown that the 520 contribution from increasing heavy events had been offset by decreasing moderate events 521 in the historical period, accounting for an insignificant rainfall trend to date. Contrasting 522 with the absence of model consensus for the historical trend, 10 of the 11 selected models 523 produce significantly more monsoon rainfall at the end of the 21st century than during any 524 of the 50-year historical period. The remaining model (FIO-ESM) does not show any trend. 525 The simulated increase in land-based rainfall ranges from 6 to 18%, except for FGOALS-s2 526 that produces 46% more rainfall at the end of the 21^{st} century (Fig. 11-c). The increase is 527 in the range 7-15% when considering the whole South Asia domain, except for FGOALS-s2 528 that produces 27% more monsoon rainfall after 2050 than in the 19th century (Fig. 11-d). 529

⁵³⁰ Finally, we investigate potential trends in the amplitude of the interannual variability of

the summer monsoon rainfall. This analysis is done in a similar way as for the mean, but 531 using 90% confidence intervals based on the χ^2 statistics (suitable for tests on standard devia-532 tions, Von Storch and Zwiers, 2002). Only FGOALS-s2 produces a strengthened interannual 533 variability over Australia (by 45%), the other models producing no significant trend until 534 2100 (not shown). Over the Maritime Continent (AMAR), four of the 11 selected models 535 show a significant increase of the interannual standard deviation (not shown). Interestingly, 536 two of these models (FGOALS-s2 and HadGEM2-AO) are among the few models that cap-537 ture a decreased mean AMAR monsoon rainfall at the end of the 21st century (as already 538 shown in Fig. 11-b). CCSM4 and FGOALS-s2 produce an increased interannual variability 539 for LIND (by 17 and 60% respectively) and for ISAS (by 20 and 42% respectively), and 540 CESM1-CAM5 produces a significantly increased interannual variability for ISAS (by 25%) 541 but not for LIND (not shown). The remaining models do not show a significant change in 542 the amplitude of the interannual variability. 543

544 **Discussion**

In this paper, we have adopted a large-scale point of view, which was needed to assess 60 545 CMIP models in a concise way. We have found some agreement between the selected models 546 for the Indian and Australian rainfall projections. Nonetheless, the spatial scale of concern 547 for human activities is much narrower than the large-scale used in this paper. Therefore, we 548 now assess the meaning, in term of regional climate, of the future increase in the large-scale 549 precipitation. Historical summer monsoon rainfall and its change in the rcp8.5 scenario are 550 shown on maps, in Fig. 12 and Fig. 13 for the Australian/Maritime Continent region and the 551 Indian monsoon respectively. As already seen in Fig. 11, a majority of the models produce 552

⁵⁵³ more monsoon rainfall over North Australia in the future. Shi et al. (2008) have reported a ⁵⁵⁴ larger rainfall increase in North-West Australia than in North-East Australia during the end ⁵⁵⁵ of the 20thcentury. There is however no such consensus in the selected rcp8.5 simulations, ⁵⁵⁶ some of them showing an East-West symmetry in the increase, or the opposite asymmetry ⁵⁵⁷ (Fig. 12-c).

In Austral summer, the Maritime Continent is at the intersection of three major conver-558 gence zones: the South Pacific Convergence Zone (SPCZ), the North Pacific Intertropical 559 Convergence Zone (ITCZ), and the South Indian Convergence Zone (SICZ), their location 560 being shown in Fig. 12-a. As such, the projections of monsoon rainfall over the Maritime 561 Continent are probably sensitive to the evolution of these convergence zones. First, it should 562 be noticed that some models tend to produce too much rainfall in the Western end of the 563 ITCZ as compared to the Western end of the SPCZ during the historical period (e.g. CCSM4, 564 FIO-ESM, MIROC5, NorESM-ME, in Fig. 12-b). Then, the striking thing is that the pattern 565 of projected DJFM rainfall in these convergence zones is totally different from one model 566 to another (Fig. 12-c). Moreover, the Maritime Continent is characterized by marked land-567 ocean heterogeneities, and by high and narrow mountain ranges, with the Central Range of 568 Papua-New Guinea peaking at 4884 m, and with mountain ranges peaking between 1000 and 569 3000 m in most of the Indonesian and Malaysian islands. It is worth mentioning that these 570 heterogeneities lead to a very strong uncertainty in any of the observational products. Most 571 of the selected models capture strong monsoon precipitation in Papua-New Guinea, and a 572 few of them capture relatively realistic island-related patterns in Indonesia and Malaysia 573 (CESM1-CAM5, ACCESS1-0, HadGEM2-AO in Fig. 12-b). While most of the models pro-574 duce an increase of precipitation in Papua-New Guinea during the 21st century, there is no 575

⁵⁷⁶ clear consensus across the selected models on the evolution of Indonesian and Malaysian
⁵⁷⁷ DJFM monsoon rainfall (Fig. 12-c).

As shown in section 3.3, 10 of the 11 selected models produce more summer monsoon 578 rainfall in India and South Asia during the 21st century compared to the historical period. 579 There are three regional spots that appear particularly intense in the observations (Fig. 13-580 a): the Western Ghats (South-Western part of the Indian peninsula), the Eastern coast of 581 the Bay of Bengal, and the Eastern third of the Himalaya. The majority of the selected 582 models simulate such spots in the historical simulations, but tend to produce much stronger 583 rainfall than observed in the Central part of the Himalaya (Fig. 13-b). It should be noted, 584 however, that the uncertainty in both satellite and station-based observations is very high 585 in this region of complex orography. Interestingly, all the models that produce more land-586 based rainfall in the rcp8.5 scenario have most of the rainfall increase located in the Himalaya 587 (Fig. 13-c). There is however no consensus across the selected models on how the summer 588 monsoon rainfall could vary along the Eastern coast of the Bay of Bengal. Approximately 589 half of the models produce slightly less rainfall in the Western Ghats during the 21st century 590 (e.g. MIROC5, HadGEM2-AO, CESM1-CAM5), in qualitative agreement with Rajendran 591 et al. (2012) who obtained such results from a high resolution atmospheric model. 592

593 5 Conclusion

In this paper, we have shown that a critical challenge in model rainfall assessment lies in the spread of observational data. Indeed, the mean summer monsoon rainfall and the amplitude of its interannual variability vary significantly across the observation datasets. The atmospheric reanalysis produce monsoon rainfall in the range of the observations uncertainty.

By building an envelope of the observations and reanalysis, it is possible to identify the 598 outliers, i.e. the models that are significantly different from the observations. Most of the 599 CMIP3 and CMIP5 models produce both Indian and Australian mean summer monsoon 600 rainfall reasonably close to the observations/reanalysis envelope. This is also true for the 601 amplitude of the interannual variability of the Indian and Australian summer monsoons. 602 The seasonal cycle of both the Indian and the Australian monsoons is in good agreement 603 across the observation products and reanalysis. Most of the CMIP3 and CMIP5 models 604 capture a seasonal cycle with a maximum rainfall at the right season, but the seasonal cycle 605 tends to be shorter or longer than observed in the CMIP5 simulations, and even more in the 606 CMIP3 simulations. Based on the mean monsoon rainfall, on the amplitude of its interannual 607 variability, and on the seasonal cycle, we select a subset of 19 models that statistically capture 608 the main characteristics of the monsoon, taking the observations uncertainty into account. 609 Then, we have evaluated the monsoon-ENSO and monsoon-IOD relationships in the 610

CMIP models, because ENSO and IOD are likely to change in a future climate, with 611 possible consequences for the monsoon. Because of their difference in seasonality, the 612 Australian/Maritime Continent monsoon-ENSO relationship and the Indian/South Asian 613 monsoon-ENSO relationship are affected by different kinds of biases in the CMIP models. 614 As already noted in previous studies related to the CMIP3 models, we have confirmed that 615 the intensity of the concomitant land-based Australian monsoon-ENSO relationship is cor-616 related to the intensity of simulated ENSO (this had already been noted by Cai et al., 2009) 617 and Colman et al., 2011 for the CMIP3 models). We have shown that the monsoon-ENSO 618 relationship over the Maritime continent is rather influenced by the ability of the models to 619 produce a cold anomaly in the climatological warm pool during El Niño events. In India 620

and South Asia, the monsoon-ENSO relationship strongly depends on the simulated sea-621 sonal cycle of ENSO, because El Niño or La Niña events are at their developing stage at 622 the beginning of the monsoon (whereas the Australian monsoon co-occurs with ENSO at 623 its mature stage). As the ENSO seasonal cycle is longer than observed in the CMIP sim-624 ulations (Taschetto et al., 2012), the CMIP models tend to produce monsoon rainfall that 625 is too much influenced by the tails of ENSO events from the previous year. The SON IOD 626 generally influences the monsoon rainfall over the Maritime Continent in the CMIP3 and the 627 CMIP5 simulations, and to a lesser extent over Australia. The strength of the monsoon-IOD 628 link in the models is correlated to the strength of the monsoon-ENSO link. 629

Based on these findings, we have empirically chosen a few criteria to refine the model 630 selection, towards models that do not present major biases with regards to the monsoon-631 ENSO relationship. We end up with 13 models that represent the statistical properties of 632 the Indian and Australian monsoon well and have also relatively good skills in simulating 633 ENSO-monsoon relationship. Twelve of these 13 models are from CMIP5. We have then 634 analyzed the change of monsoon rainfall in the rcp8.5 emission scenario for the 11 available 635 CMIP5 models. A large majority of these 11 models produce significantly more summer 636 monsoon rainfall in India (10/11), in the South Asia region (10/11), and in Australia (9/11)637 at the end of the 21^{st} century. Thus, the models generally produce 5 to 20% more summer 638 monsoon rainfall in 2050-2099 as compared to the pre-industrial period (and much more in 639 the FGOALS-s2 model). In India, most of the simulated increase takes place in the Himalaya. 640 By contrast, only five of the 11 models produce significantly more monsoon rainfall over the 641 Maritime Continent at the end of the 21st century. Two models (FGOALS-s2 and HadGEM2-642 AO) project slightly less monsoon rainfall over the Maritime Continent in the future, but 643

associated with a strengthened interannual variability. For the majority of the models, there
is no significant change in the amplitude of the interannual monsoon rainfall variability.
Considering maps of projected rainfall patterns, we find no consistency between the selected
models over the Maritime Continent.

Our concluding remark is that the best CMIP5 models have stronger skills than the best CMIP3 models, but the best models are still unable to resolve the complexity of the Maritime Continent. This leads to the absence of model consensus concerning the future monsoon rainfall in this region. It is likely that high-resolution modeling is needed to simulate the climate of this region, due to complex land/sea distribution and to complex orography and bathymetry.

654 Acknowledgment

This study was conducted in the context of the ARC project DP110100601. KA received 655 supports from IITM and MoES-NERC. We acknowledge the World Climate Research Pro-656 gramme's Working Group on Coupled Modelling, which is responsible for CMIP, and we 657 thank the climate modeling groups (Tab. 3, 4) for producing and making available their 658 model output. The U.S. Department of Energy's Program for Climate Model Diagnosis 659 and Intercomparison (PCMDI) provided coordinating support and led development of soft-660 ware infrastructure in partnership with the Global Organization for Earth System Science 661 Portals. We thank the Australian National Computational Infrastructure (NCI) for help in 662 the download process. We acknowledge all the Institutions in Tab. 2 for having made their 663 observations and reanalysis accessible to us. 664

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Index	Long name	Var.	References
LIND	land-only Indian monsoon	precip	
ISAS	South Asia / Indian monsoon	precip	Meehl and Arblaster, 2002
LAUS	land-only Australian monsoon	precip	
AMAR	Maritime Continent / Australian monsoon	precip	Meehl and Arblaster, 2002
NINO34	ENSO central index	SST	
DMI	(Indian) Dipole Mode Index	SST	Saji and Yamagata, 2003
IOBW	Indian Ocean Basin Wide Warming index	SST	Taschetto et al., 2011

Table 1: Indices used in this paper (see Fig. 1).

Acronym	Ι	Institute	Spatial	Start	Resol	References
	D		coverage	date		
CMAP	α	UCAR/NCAR/CISL/DSS	Global	1979	2.5°	Xie and Arkin, 1997
GPCP	β	NOAA/OAR/ESRL PSD,	Global	1979	2.5°	Adler et al., 2003
GPCC	γ	Boulder, CO, USA	Global land	1901	0.5°	Rudolf et al., 2011
AWAP	δ	BOM, Australia	Austr. land	1900	0.25°	Jones et al., 2009
APHRODITE	δ	ERTDF, Japan	S-E Asia land	1951	0.25°	Yatagai et al., 2012
TRMM-3B42 v 6	ϵ	NASA/GIES/DISC,	$50^{\rm o}{\rm S}$ -50°N	1998	0.25°	Adler et al., 2000
TRMM-3B43 v 6	ζ	USA	$50^{\rm o}{\rm S}$ -50°N	1998	0.25°	Adler et al., 2000
HadISST	η	Met Office,	Global	1870	1.0°	Rayner et al., 2003
HadSST2	θ	Hadley Centre, UK	Global	1850	5.0°	Rayner et al., 2006
NCEP-NCAR I	λ	NOAA/OAR/ESRL PSD,	Global	1948	2.5°	Kalnay et al., 1996
NCEP-DOE II	μ	Boulder, CO,	Global	1979	2.5°	Kalnay et al., 1996
NCEP-CFSR	π	USA	Global	1979	0.5°	Saha et al., 2010
ERA-40	ρ	ECMWF, UK	Global	1957	2.5°	Dee et al., 2011
ERAinterim	au	ECMWF, UK	Global	1979	0.7°	Dee et al., 2011
JRA-25	ψ	JMA/CRIEPI, Japan	Global	1979	2.5°	Onogi et al., 2007
MERRA	σ	NASA	Global	1979	0.5°	Rienecker et al., 2011

Table 2: Observations (precipitation in upper part, SST in the middle part) and atmospheric reanalysis (lower part) used in this paper. If "land" is not mentioned, precipitation datasets cover both land and ocean. Most of the datasets cover up to the recent years (around 2010), except ERA40 that was stopped in 2002 and APHRODITE that is only available until 2007. The resolution mentioned here is the one of the gridded dataset, even if most of the reanalysis are produced using spectral models.

			Ocean	Atmos.	Ens	
Model	Ι	Institute	horizontal	horiz.	O/A	References
	D		resolution	resol.		
bccr-bcm2-0	a	BCCR, Norway	1.0×1.0 (1.0)	2.8×2.8	1/1	Furevik et al., 2003
cccma-cgcm3-1	b	CCCMA, BC,	1.9×1.9 (1.9)	$3.7{\times}3.7$	2/5	Kim et al., 2002
cccma-cgcm3-1-t63	c	Canada	$1.4 \times 0.9 \ (0.9)$	$2.8{\times}2.8$	1/1	Kim et al., 2002
cnrm-cm3	d	CNRM, France	$2.0 \times 1.0 (1.0)$	$2.8{\times}2.8$	1/1	Salas-Mélia et al., 2005
csiro-mk3-0	е	CSIRO,	$1.9{ imes}0.9~(0.9)$	$1.9{ imes}1.9$	1/2	Gordon et al., 2002
csiro-mk3-5	f	Australia	$1.9{ imes}0.9~(0.9)$	$1.9{ imes}1.9$	1/1	Gordon et al., 2002
gfdl-cm2-0	g	NOAA, GFDL,	$1.0 \times 1.0 \ (0.4)$	2.5×2.0	1/1	Delworth et al., 2006
gfdl-cm2-1	h	USA	$1.0 \times 1.0 \ (0.4)$	2.5×2.0	1/3	Delworth et al., 2006
giss-aom	i	NASA/GISS,	4.0×3.0 (3.0)	4.0×3.0	1/2	Lucarini and Russell, 2002
giss-model-e-h	j	USA	$1.0 \times 1.0 (1.0)$	$5.0{\times}4.0$	4/5	Schmidt et al., 2006
giss-model-e-r	k	" " "	5.0×4.0 (4.0)	$5.0{\times}4.0$	9/9	Schmidt et al., 2006
iap-fgoals1-0-g	1	IAP, China	$1.0 \times 1.0 (1.0)$	$2.8{\times}2.8$	3/3	Yongqiang et al., 2004
ingv-echam4	m	INGV, Italy	$1.0 \times 1.0 (1.0)$	1.1×1.1	1/1	Gualdi et al., 2008
inmcm3-0	n	INM, Russia	2.5×2.0 (2.0)	$5.0{\times}4.0$	1/1	Volodin and Diansky, 2004
ipsl-cm4	0	IPSL, France	$2.0 \times 1.0 (1.0)$	$3.7{\times}2.5$	1/1	Marti et al., 2005
miroc3-2-hires	р	CCSR,	$1.2 \times 0.6 \ (0.6)$	1.1×1.1	1/1	K-1 Developers, 2004
miroc3-2-medres	q	Japan	$1.4 \times 0.9 \ (0.6)$	$2.8{\times}2.8$	1/3	K-1 Developers, 2004
miub-echo-g	r	MIUB, Germany	$2.8 \times 2.3 \ (0.5)$	$3.7{\times}3.7$	2/5	Min and Hense, 2006
		& Korea				
mpi-echam5	\mathbf{S}	MPI-M, Germany	$1.0 \times 1.0 (1.0)$	$1.9{ imes}1.9$	1/3	Jungclaus et al., 2006
mri-cgcm2-3-2a	\mathbf{t}	MRI, Japan	$2.5 \times 2.0 \ (0.5)$	$2.8{\times}2.8$	5/5	Yukimoto et al., 2001
ncar-ccsm3-0	u	NCAR, CO, USA	$1.1 \times 0.5 \ (0.3)$	1.4×1.4	2/8	Collins et al., 2006
ncar-pcm1	v	NCAR, CO, USA	1.0×1.0 (1.0)	$2.8{\times}2.8$	3/4	Washington et al., 2000
ukmo-hadcm3	W	MOHC, UK	$1.2 \times 1.2 (1.2)$	$3.8{\times}2.5$	1/2	Gordon et al., 2000
ukmo-hadgem1	х	MOHC, UK	$1.0 \times 1.0 \ (0.4)$	$1.9{ imes}1.2$	1/2	Johns et al., 2004

Table 3: CMIP3 model names; ID for this paper; name of providing institutes; ocean mean zonal resolution (at the equator in °E) × mean 25°N -35°N resolution in latitude (in °), and mean equatorial refinement in brackets (5°S -5°N); atmospheric output resolution (in °E × °N); the number of ensemble members for the Ocean/Atmosphere components is shown in the Ens O/A column.

			Ocean	Atmos.	Ens	
Model	Ι	Institute	horizontal	horiz.	O/A	References
	D		resolution	resol.		
ACCESS-1.0	А	CSIRO-BOM,	$1.0 \times 1.0 \ (0.3)$	$1.9{\times}1.2$	1/1	BOM, 2010
ACCESS-1.3	В	Australia	$1.0 \times 1.0 \ (0.3)$	$1.9{ imes}1.2$	1/1	BOM, 2010
BCC-CSM1-1	\mathbf{C}	BCC, CMA, China	$1.0 \times 1.0 \ (0.3)$	2.8×2.8	3/3	
CanESM2	D	CCCMA, Canada	$1.4 \times 0.9 \ (0.9)$	$2.8{\times}2.8$	5/5	
CESM1-CAM5	Е	NSF-DOE-	$1.1 \times 0.6 \ (0.3)$	$1.2{ imes}0.9$	2/3	Vertenstein et al., 2012
CESM1-FASTCHEM	\mathbf{F}	-NCAR,	$1.1 \times 0.6 \ (0.3)$	$1.2{ imes}0.9$	3/3	Vertenstein et al., 2012
CESM1-WACCM	G	USA	$1.1 \times 0.6 \ (0.3)$	2.5×1.9	1/1	Vertenstein et al., 2012
CCSM4	Η	NCAR, CO, USA	$1.1 \times 0.6 \ (0.3)$	$1.2{\times}0.9$	1/6	Gent et al., 2011
CMCC-CM	Ι	CMCC, Italia	$2.0 \times 1.9 \ (0.6)$	$0.7{ imes}0.7$	1/1	Scoccimarro et al., 2011
CNRM-CM5	J	CNRM-CERFACS,	$1.0 \times 0.8 \ (0.3)$	1.4×1.4	10	Voldoire et al., 2012
		France			/10	
CSIRO-Mk3-6-0	Κ	CSIRO-QCCCE,	$1.9{ imes}0.9~(0.9)$	$1.9{ imes}1.9$	10	Rotstayn et al., 2012
		Australia			/10	Rotstayn et al., 2010
EC-EARTH	\mathbf{L}	EC-EARTH, Europe	$1.0 \times 0.8 \ (0.3)$	1.1×1.1	0/2	Hazeleger et al., 2010
FGOALS-g2	Μ	LASG-CESS, China	$1.0 \times 1.0 \ (0.5)$	2.8×2.8	2/3	Yongqiang et al., 2004
FGOALS-s2	Ν	LASG-IAP, China	$1.0 \times 1.0 \ (0.5)$	$2.8{ imes}1.7$	2/3	Yongqiang et al., 2004
FIO-ESM	0	FIO, SOA, China	$1.1 \times 0.6 \ (0.3)$	2.8×2.8	1/1	
GFDL-CM3	Р	NOAA	$1.0 \times 1.0 \ (0.4)$	2.5×2.0	1/5	Donner et al., 2011
GFDL-ESM2G	Q	-GFDL,	$1.0 \times 1.0 \ (0.4)$	2.5×2.0	1/3	Donner et al., 2011
GFDL-ESM2M	R	USA	$1.0 \times 1.0 \ (0.4)$	2.5×2.0	1/1	Donner et al., 2011
GISS-E2-H	\mathbf{S}	NASA/GISS,	2.5×2.0 (2.0)	2.5×2.0	5/5	Schmidt et al., 2006
GISS-E2-R	Т	NY, USA	2.5×2.0 (2.0)	2.5×2.0	5/4	Schmidt et al., 2006
HadCM3	U	MOHC, UK	$1.2 \times 1.2 (1.2)$	$3.7{\times}2.5$	9/4	Collins et al., 2001
HadGEM2-AO	V	NIMR-KMA,Korea	$1.0 \times 1.0 \ (0.4)$	$1.9{\times}1.2$	1/1	Martin et al., 2011
HadGEM2-CC	W	MOHC, UK	$1.0 \times 1.0 \ (0.4)$	$1.9{ imes}1.2$	2/3	Martin et al., 2011
HadGEM2-ES	Х	MOHC, UK	$1.0 \times 1.0 \ (0.4)$	$1.9{ imes}1.2$	2/3	Collins et al., 2011
INMCM4	Υ	INM, Russia	$0.8 \times 0.4 \ (0.4)$	2.0×1.5	1/1	Volodin et al., 2010
IPSL-CM5A-LR	Ζ	IPSL, France	$2.0 \times 1.9 \ (0.6)$	$3.7{\times}1.9$	4/4	Dufresne et al., 2012
IPSL-CM5B-LR	Г	IPSL, France	$2.0 \times 1.9 \ (0.6)$	$3.7{\times}1.9$	1/1	Dufresne et al., 2012
IPSL-CM5A-MR	Δ	IPSL, France	$1.6 \times 1.4 \ (0.6)$	2.5×1.3	1/1	Dufresne et al., 2012
MIROC5	Π	AORI-NIES-	$1.6 \times 1.4 \ (0.6)$	1.4×1.4	3/3	Watanabe et al., 2010
MIROC-ESM	Σ	-JAMSTEC, Japan	$1.4 \times 0.9 \ (0.6)$	2.8×2.8	3/3	Watanabe et al., 2011
MPI-ESM-LR	Ω	MPI-N, Germany	$1.5 \times 1.5 (1.5)$	1.9×1.9	3/3	Raddatz et al., 2007
MPI-ESM-MR	0	MPI-N, Germany	0.4×0.4 (0.4)	1.9×1.9	3/3	Raddatz et al., 2007
MRI-CGCM3	#	MRI, Japan	$1.0 \times 0.5 \ (0.5)$	1.1×1.1	3/3	Yukimoto et al., 2001
NorESM1-M	\$	NCC, Norway	$1.1 \times 0.6 \ (0.3)$	2.5×1.9	3/3	
NorESM1-ME	&	NCC, Norway	$1.1 \times 0.6 \ (0.3)$	2.5×1.9	1/1	

Table 4: CMIP5 model names; ID for this paper; name of providing institutes; ocean mean zonal resolution (at the equator in °E) × mean 25°N -35°N resolution in latitude (in °), and mean equatorial refinement in brackets (5°S -5°N); atmospheric output resolution (in °E × °N); the number of ensemble members for the Ocean/Atmosphere components is shown in the Ens O/A column.

model	ID	ensemble	AMAR	LAUS	LIND
name		members	-ENSO	-ENSO	-ENSO
gfdl-cm2-0	g	—			\checkmark
ACCESS-1.0	А	1	\checkmark	\checkmark	\checkmark
CanESM2	D	5		\checkmark	\checkmark
CESM1-CAM5	Ε	3		\checkmark	\checkmark
CESM1-FASTCHEM	\mathbf{F}	0	\checkmark		\checkmark
CCSM4	Η	6		\checkmark	\checkmark
CNRM-CM5	J	5		\checkmark	\checkmark
FGOALS-g2	Μ	_	\checkmark		
FGOALS-s2	Ν	3			\checkmark
FIO-ESM	Ο	3			\checkmark
GFDL-CM3	Р	_			\checkmark
GFDL- $ESM2G$	Q	_			\checkmark
HadCM3	U	_			\checkmark
HadGEM2-AO	V	1			\checkmark
MIROC5	П	1			\checkmark
MPI-ESM-LR	Ω	_		\checkmark	\checkmark
MPI-ESM-MR	0	—		\checkmark	\checkmark
NorESM1-M	\$	1	\checkmark	\checkmark	\checkmark
NorESM1-ME	&	1	\checkmark	\checkmark	\checkmark

Table 5: List of the CMIP models showing the best skills in term of Indo-Australian monsoon statistical properties. The number of ensemble members used for the assessment of rcp8.5 projections is shown in the third column. The symbol $\sqrt{}$ indicates that the specified monsoon-ENSO relationship is well captured, based on criteria described in section 3.2. The symbol \frown indicates a failure in the specified monsoon-ENSO relationship. The models that are removed due to deficiency in the simulated ENSO-monsoon relationship are in italic, the other are in bold.

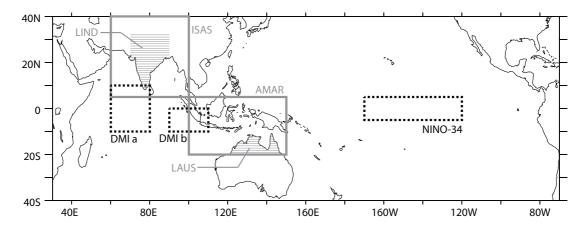


Figure 1: Boxes used to compute indices defined in Tab. 1 (DMI is calculated as DMIa-DMIb), with the shaded areas showing the land-based indices.

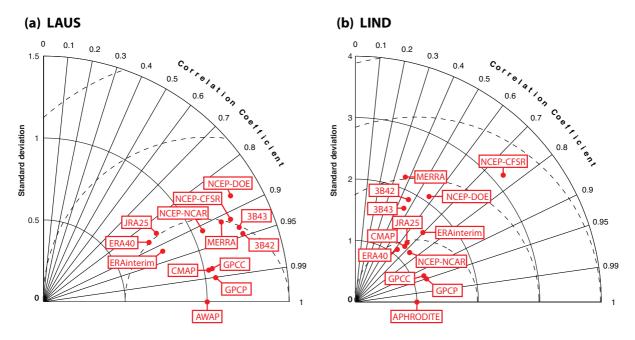


Figure 2: (a) Taylor (2001) diagram for the DJFM Australian monsoon rainfall index (LAUS). (b) same for the JJAS Indian monsoon rainfall index (LIND). One standard deviation unit on the diagram is one standard deviation of AWAP and APHRODITE in Australia and India respectively. Each dataset is compared to AWAP/APHRODITE over the common period (*e.g.* 1948-2009 for NCEP-NCAR but 1998-2009 for TRMM-3B43 in Australia). The distance from AWAP/APHRODITE represents the centered root mean square error as compared to AWAP/APHRODITE (dashed lines, in AWAP/APHRODITE standard deviation units).

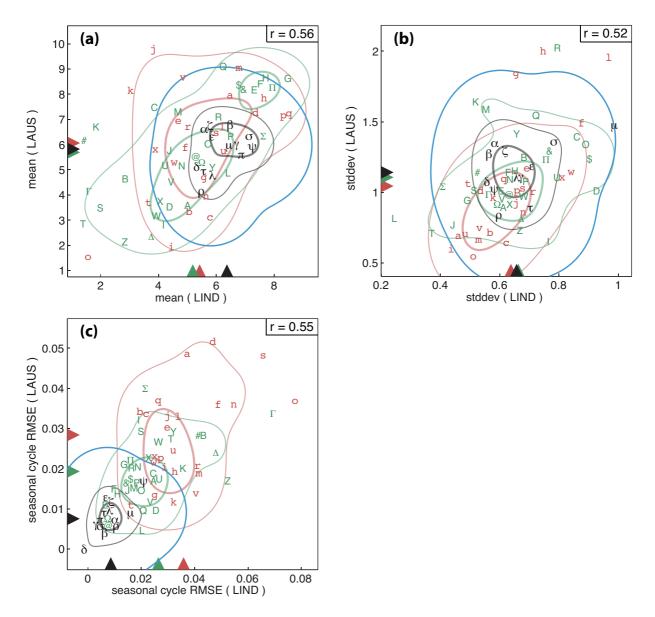


Figure 3: (a) mean DJFM Australian monsoon rainfall (LAUS) as a function of mean JJAS Indian monsoon rainfall (LIND). (b) Same as (a) but for interannual standard deviations of summer-months-averaged rainfall instead of means. (c) RMSE of normalized LAUS seasonal cycles versus RMSE of normalized LIND seasonal cycles, the RMSE being calculated with respect to AWAP in Australia and to APHRODITE in India (see detailed seasonal cycles in Fig. 4). Each letter or symbol refers to a model/dataset from CMIP5 (green, Tab. 4), CMIP3 (red, Tab. 3), or reanalysis/observations (black, Tab. 2). Triangles show the multi-model mean. Units in (a) and (b) are mm/day, while (c) has no units. The number r (upper right of each panel) is the correlation coefficient of the X - Y scatter plot, for CMIP3 and CMIP5 considered together (without observations and reanalysis). The PDF contours are estimated from the sum up of Gaussian functions attributed to each model point. The standard deviation of each individual Gaussian function is chosen as $3s/\sqrt{N}$ in each direction, where s is the standard deviation of one group (CMIP3, CMIP5, or observations/reanalysis), and N the number of elements within the group (such a standard deviation for the Gaussian function enables to fill the average distance between two neighbor points among N points normally distributed). Thick (thin) black, red, and green contours enclose 25% (75%) of PDF integrative. The blue contour encloses 99.9% of the observations/reanalysis PDF integrative in (a) and (b), and 99.999% in (c). In (b), NCEP-DOE-II (represented by μ) and NCEP-CFSR (out of the figure area, standard deviation of 1.5 and 2.0 mm/day for LAUS and LIND respectively) are not considered in the PDF computation. In (c), JRA25 is not considered in the PDF computation.

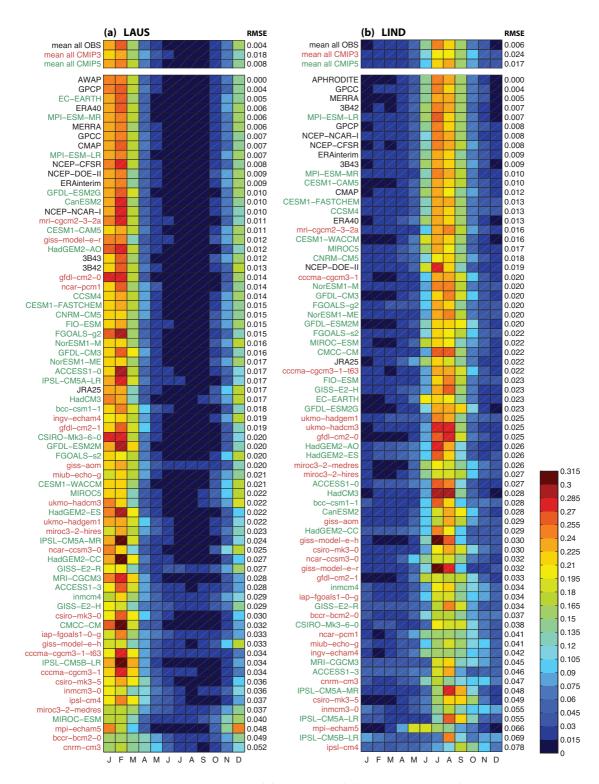


Figure 4: Normalized seasonal cycle of LAUS (a) and LIND (b), for observations/reanalysis (black), CMIP3 (red) and CMIP5 (green), and sorted by increasing RMSEs (shown on the right of each panel, and computed with regards to AWAP in Australia and to APHRODITE in India).

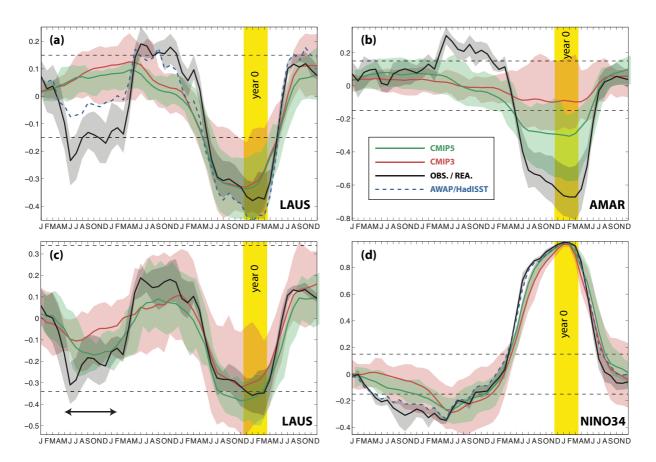


Figure 5: (a) Lag correlation between LAUS averaged in DJFM of year 0 and monthly NINO34 values (months on the X-axis). Thick lines are the means over the observations/reanalysis (black), CMIP3 (red), and CMIP5 (green). Semi-transparent areas show the upper and lower quartiles. The dashed blue thick line in (a) and (d) represents AWAP-HadISST. The yellow area indicates the reference time (t=0), and its width shows the DJFM months over which each index is averaged. (b) Same as (a) but for AMAR instead of LAUS. (c) Same as (a), but selecting the 30 year period of each model (over the 150 years of each member and among the ensemble members) that gives the strongest anti-correlation over the period shown by the black double arrow. (d) Same as (a) but for NINO34 instead of LAUS.

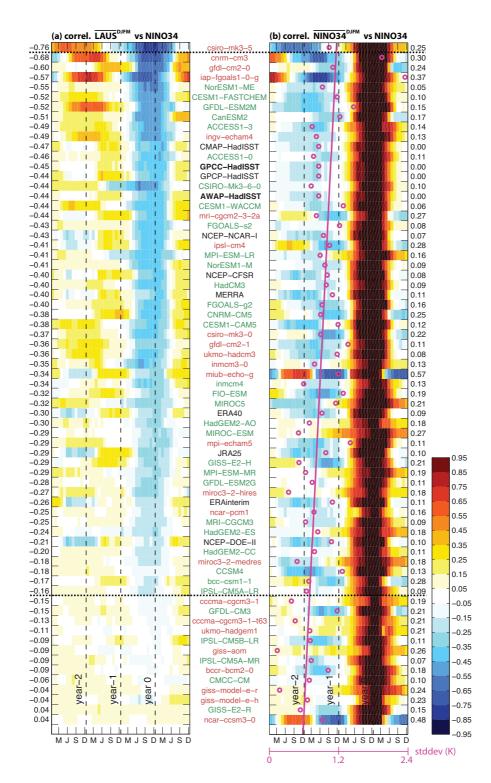


Figure 6: (a) Lag correlation between LAUS averaged in DJFM of year 0 and monthly NINO34 values (months on the X-axis, M for March and J for June) for observations/reanalysis (black names), CMIP3 (red), and CMIP5 (green) ranked by increasing correlation in year0 DJFM. (b) Same as (a) but for correlation between NINO34 averaged in DJFM and lagged monthly NINO34 values. Pink circles indicate the standard deviation of NINO34 produced by each model. The pink line is the least mean square linear fit of these circles (correlation=0.45).

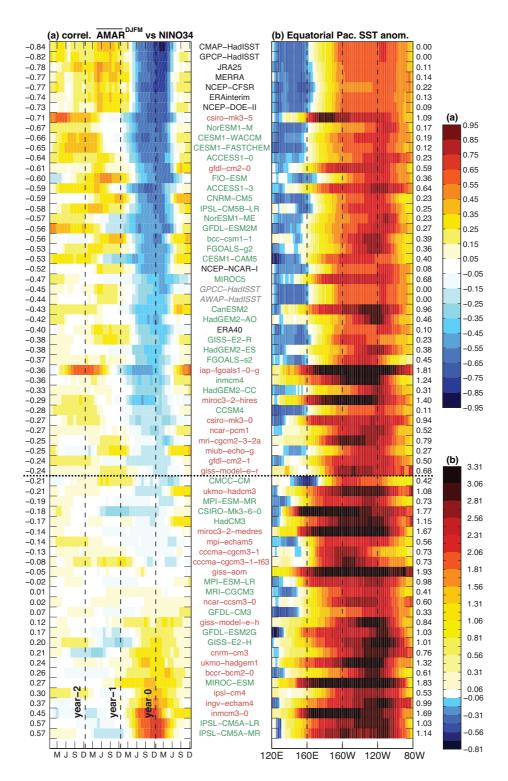


Figure 7: (a) Lag correlation between AMAR averaged in DJFM of year 0 and monthly NINO34 values (months on the X-axis, M for March and J for June) for observations/reanalysis (black names), CMIP3 (red), and CMIP5 (green) ranked by increasing correlation in year0 DJFM (indicated on the left hand side). Names of land-only data are in gray. (b) Composite of equatorial Pacific SST anomalies (5°N -5°S average, in K) for NINO34 greater than 1 standard deviation. The RMSE with regards to HadISST is indicated on the right hand side.

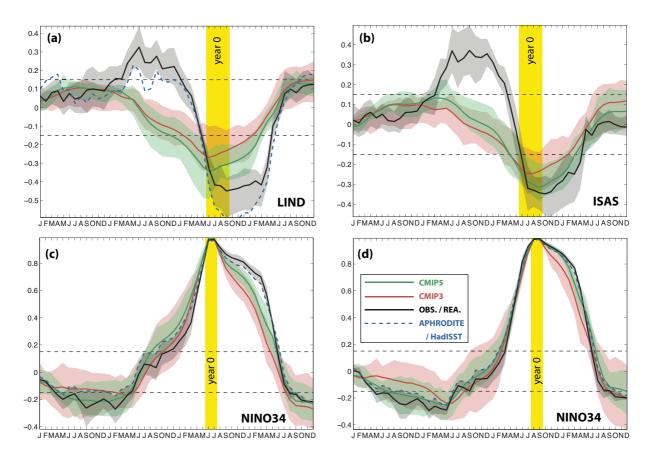


Figure 8: (a) Lag correlation between LIND averaged in JJAS of year 0 and monthly NINO34 values (months on the X-axis). Thick lines are the means over the observations/reanalysis (black), CMIP3 (red), and CMIP5 (green). Semi-transparent areas show the upper and lower quartiles. The dashed blue thick line represents APHRODITE-HadISST. The yellow area indicates the reference time (t=0), and its width shows the JJAS months over which each index is averaged. The black dashed lines represent the 90% significance of correlation coefficients for a single time-series of 150 years (see caption of Fig. 5). (b) Same as (a) but for ISAS instead of LIND. (c) Same as (a) but for June-July NINO34 instead of JJAS LIND. (d) Same as (c) but for August-September NINO34.

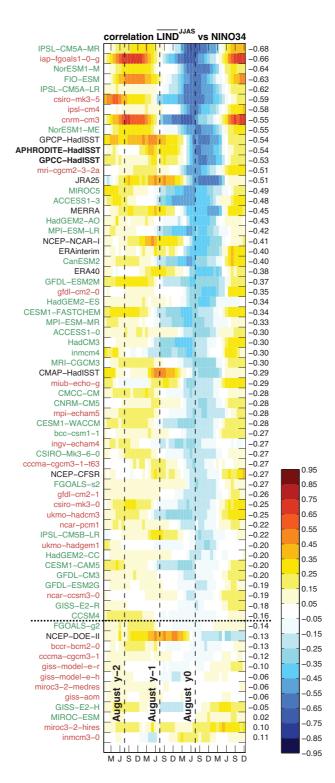


Figure 9: Lag correlation between LIND averaged in JJAS of year 0 and monthly NINO34 values (months on the X-axis, M for March and J for June) for observations/reanalysis (black names), CMIP3 (red), and CMIP5 (green) ranked by increasing correlation in year0 JJAS.

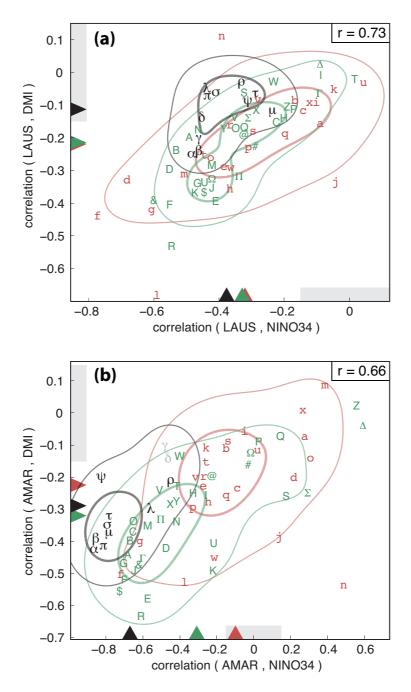


Figure 10: (a) Scatter plot of the correlation between SON DMI and DJFM LAUS (Y-axis) as a function of the correlation between DJFM NINO34 and DJFM LAUS (X-axis). Envelopes are PDF contours (see Fig. 3). Multi-model means are shown by triangles. Gray bars along the axis show correlation values that are below the 90% significance level for a 150-year time series. (b) Same but with AMAR instead of LAUS, with land-based observations represented by gray letters.

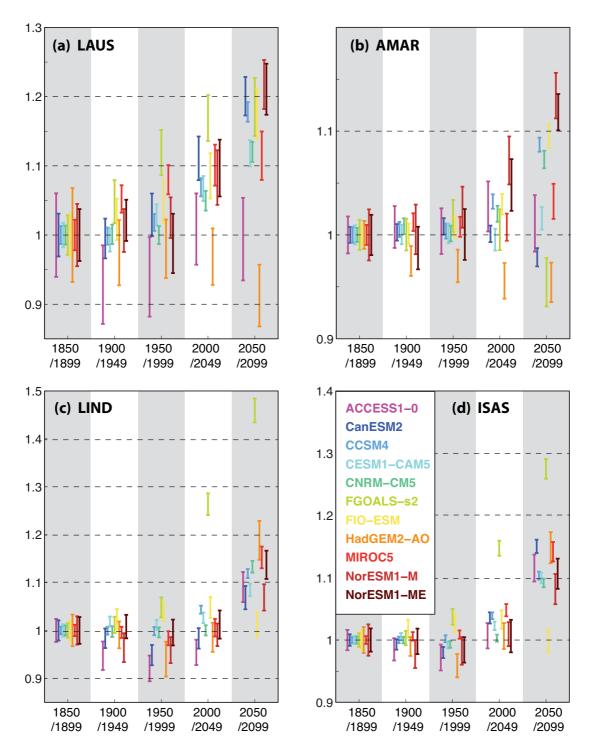


Figure 11: Summer monsoon rainfall averaged over 50-year periods and divided by the 1850-1899 mean for LAUS (a), AMAR (b), LIND (c), and ISAS (d). Error bars show the confidence interval at the 90% level (the uncertainty of the denominator not taken into account, so that bars have to be compared to each others rather than to unity).

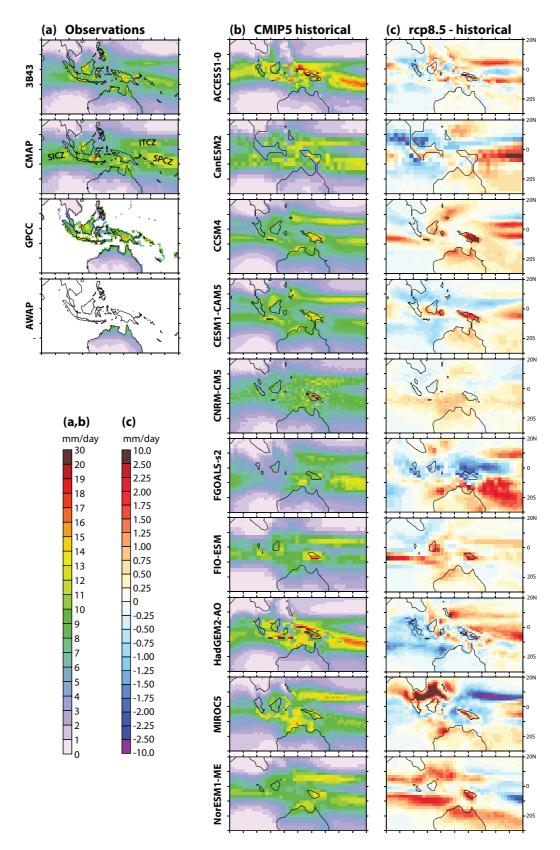


Figure 12: DJFM rainfall in the observations (a) and in the historical CMIP5 simulations (b), and difference between the 2006-2100 mean rainfall from rcp8.5 experiments and the 1850-2005 mean rainfall (c). Maps from NorESM1-M are not shown since they are quite similar to the map from NorESM1-ME.

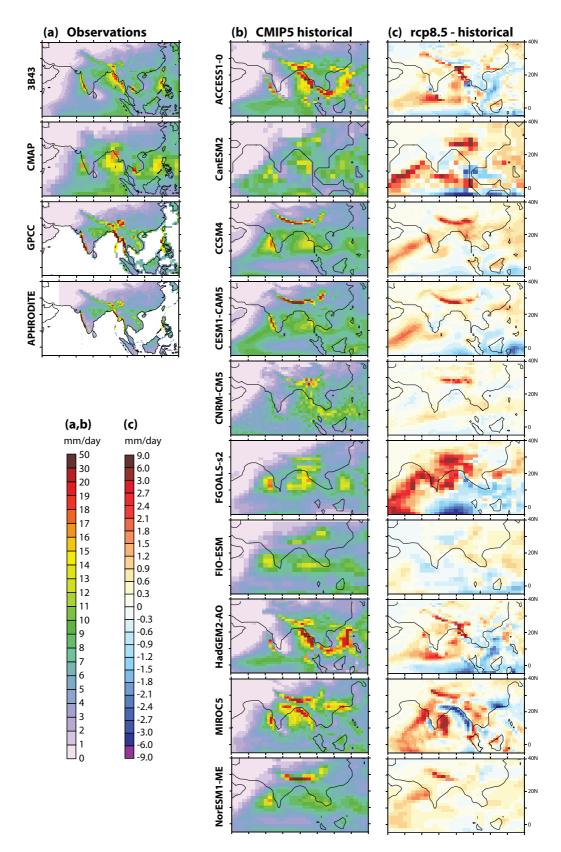


Figure 13: JJAS rainfall in the observations (a) and in the historical CMIP5 simulations (b), and difference between the 2006-2100 mean rainfall from rcp8.5 experiments and the 1850-2005 mean rainfall (c). Maps from NorESM1-M are not shown since they are quite similar to the map from NorESM1-ME.